#### Chiral Shock Waves

#### Srimoyee Sen, University of Arizona

In collaboration with Naoki Yamamoto (Keio University)

ArXiv: 1609.07030v2

#### Outline

- Chiral transport phenomenon.
- What is a shock wave ?
- Shock wave in nonchiral matter.
- Shockwave in chiral matter.
- Example and implications.

# Ideal and dissipative hydrodynamics

• Energy momentum conservation and particle number conservation:  $\partial_{\mu}T^{\mu\nu}=0,$ 

$$\partial_{\mu}j^{\mu}=0.$$

Where 
$$T^{\mu\nu}=hu^{\mu}u^{\nu}-pg^{\mu\nu}$$
  $j^{\mu}=nu^{\mu}$   $h=\epsilon+p$   $u^{\mu}=\gamma(1,{m v})$ 

 Dissipative processes require additional terms in the conserved quantities – to be constrained by the second law of thermodynamics.

# Background field and anomaly

- Consider massless fermions of single chirality.
- In the presence of a background electromagnetic field

$$\partial_{\mu}T^{\mu\nu} = F^{\nu\lambda}j_{\lambda} \,, \quad \partial_{\mu}j^{\mu} = CE^{\mu}B_{\mu}$$

 External fields add further dissipative components to conserved currents where as anomaly adds nondissipative terms.

## Chiral Transport

 Ignoring dissipation the constitutive relations become in Landau frame:

$$j^{\mu} = nu^{\mu} + \xi \omega^{\mu} + \xi_B B^{\mu}$$





#### chiral vortical effect chiral magnetic effect

$$\xi = \frac{C}{2}\mu^{2} \left( 1 - \frac{2}{3} \frac{n\mu}{\epsilon + P} \right) + \frac{D}{2}T^{2} \left( 1 - \frac{2n\mu}{\epsilon + P} \right)$$

$$\xi_{B} = C\mu \left( 1 - \frac{1}{2} \frac{n\mu}{\epsilon + P} \right) - \frac{D}{2} \frac{nT^{2}}{\epsilon + P}$$

$$\omega^{\mu} = \frac{1}{2} \epsilon^{\mu\nu\lambda\rho} u_{\nu} \partial_{\lambda} u_{\rho} \quad B^{\mu} = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} u_{\nu} F_{\alpha\beta}$$

## Fluid flow (sonic)

- Effects of compressibility for high velocity fluid flow important.
- Two types: subsonic and supersonic flow.
- Supersonic flow of two types:
  - Steady continuous flow.
  - Surface of discontinuity → Shock waves.

### Surface of discontinuity

- Velocity, pressure, density and temperature can be discontinuous across a surface perpendicular to the flow.
- Certain boundary conditions must be satisfied at this surface.
- Quantities like mass flux and energy flux should remain continuous across the surface for nonrelativistic and relativistic systems respectively.

## Nonchiral relativistic Shock Waves

- Shockwave traveling along 'x' axis.
- The region behind and ahead of the shock wave front denoted by 1 and 2.
- Impose continuity in particle number flux, energy flux etc.

$$j_1^x = j_2^x$$
  $T_1^{xx} = T_2^{xx}$   $T_1^{0x} = T_2^{0x}$   
 $T_1^{yx} = T_2^{yx}$   $T_1^{zx} = T_2^{zx}$ 

#### Pressure-volume relation

$$v_1 = \sqrt{\frac{(p_2 - p_1)(\epsilon_2 + p_1)}{(\epsilon_2 - \epsilon_1)(\epsilon_1 + p_2)}}$$

$$v_2 = \sqrt{\frac{(p_2 - p_1)(\epsilon_1 + p_2)}{(\epsilon_2 - \epsilon_1)(\epsilon_2 + p_1)}}$$

The adiabatic: 
$$\frac{v_1}{V_1\sqrt{1-v_1^2}} - \frac{v_2}{V_2\sqrt{1-v_2^2}} = 0$$

#### Weak shock waves

$$\epsilon_2 \to \epsilon_1 \qquad p_2 \to p_1 \quad \Delta \epsilon \equiv \epsilon_2 - \epsilon_1 \quad \Delta p \equiv p_2 - p_1$$

$$v_1^2 = (c_{s1})^2 \left( 1 + \frac{\Delta \epsilon}{h_1} (1 - c_{s1}^2) + \dots \right)$$

$$v_2^2 = (c_{s2})^2 \left( 1 - \frac{\Delta \epsilon}{h_1} (1 - c_{s2}^2) + \dots \right)$$

$$c_s^2 = \frac{dp}{d\epsilon}$$

## Compression and rarefaction shock waves

$$\epsilon_2 > \epsilon_1$$



$$\epsilon_2 > \epsilon_1 \qquad (v_1)^2 > (v_2)^2$$

$$p_2 > p_1$$

$$p_2>p_1$$
 follows from:  $c_{\rm s}^2=\lim_{2\to 1}\frac{p_2-p_1}{\epsilon_2-\epsilon_1}>0$ 

----- compression

$$\epsilon_2 < \epsilon_1$$



$$\epsilon_2 < \epsilon_1 \qquad \qquad (v_1)^2 < (v_2)^2$$

$$p_1 > p_2$$

$$p_1>p_2$$
 follows from:  $c_{\rm s}^2=\lim_{2\to 1}\frac{p_2-p_1}{\epsilon_2-\epsilon_1}>0$ 

---- rarefaction

Define

$$\Delta H = H_2 - H_1 \ \Delta S = S_2 - S_1$$
$$\Delta V = V_2 - V_1$$

Expand the pressure-volume relation using:

$$\Delta H = T\Delta S + V_1 \Delta p + \frac{1}{2} \frac{\partial V}{\partial p} \Big|_{1} (\Delta p)^2 + \frac{1}{6} \frac{\partial^2 V}{\partial p^2} \Big|_{1} (\Delta p)^3 + \cdots$$

$$\Delta V = \frac{\partial V}{\partial p} \Big|_{1} \Delta p + \frac{1}{2} \frac{\partial^2 V}{\partial p^2} \Big|_{1} (\Delta p)^2 + \frac{1}{6} \frac{\partial^3 V}{\partial p^3} \Big|_{1} (\Delta p)^3$$

$$+ \frac{\partial V}{\partial S} \Big|_{1} \Delta S + \cdots$$

## Entropy discontinuity and shockwaves

$$\Delta S = \frac{1}{12H_1T} \frac{\partial^2 (HV)}{\partial p^2} \bigg|_1 (\Delta p)^3 + O((\Delta p)^4)$$

For any realistic equation of state:  $\partial^2(HV)/\partial p^2>0$ 

Second law of thermodyncamics:

$$S_2 > S_1 \longrightarrow p_2 > p_1 \longrightarrow$$

$$p_1 > p_2 \longrightarrow$$

Hence, only compression shock waves allowed!

#### Chiral shock waves

- Does the pressure entropy discontinuity depend on chiral transport? How?
- To answer choose a particular limit.
- Set  $B^{\mu}$  to 0 such that, but  $\omega^{\mu}$  is finite

$$j^{\mu} = nu^{\mu} + \xi \omega^{\mu}$$

• Fermions with chemical potential  $\mu$  and temperature T such that  $T/\mu \ll 1$ 

#### Chiral shock waves

- Shock wave front traveling along 'x'.
- Vorticity chosen to be along 'x',

$$\omega_x = \omega, \omega_y = \omega_z = 0$$

- Hydrodynamics makes sense when  $\omega \ll \mu$
- Back of the wave-front denoted by '1', Front denoted by '2'.

#### Chiral shock waves

- Due to this vorticity, we cannot go to a frame with  $v_1^y=v_2^y=v_1^z=v_2^z=0$  everywhere.
- Consider the regime  $|\omega|\rho\ll 1$  where the distance from axis:  $\rho=\sqrt{y^2+z^2}$
- In that case the speed perpendicular to the direction of vorticity

$$v_{\perp} = \omega \rho (1 - v_x^2) + O((\omega \rho)^2)$$

## Continuity equation

From the continuity equation we have to have

$$v_1^y=v_2^y$$
 and  $v_1^z=v_2^z$  or  $v_1^\perp=v_2^\perp$ 

This implies

$$\omega_1 \left( 1 - (v_1^x)^2 \right) = \omega_2 \left( 1 - (v_2^x)^2 \right)$$

• In the limit  $|\omega|\rho\ll 1$  the expressions for  $v_1^2$  and  $v_2^2$  are given by their nonchiral version.

#### Pressure-volume relation

$$\frac{v_1}{V_1\sqrt{1-v_1^2}} - \frac{v_2}{V_2\sqrt{1-v_2^2}} = -(\xi_1\omega_1 - \xi_2\omega_2)$$

- We know the expansion of the LHS in terms of  $\Delta S$  and  $\Delta p$ .
- The RHS is a function of  $\mu_1, \mu_2, T_1, T_2$  as well.
- We do not know the expansion of  $\Delta \mu \equiv \mu_2 \mu_1$  and  $\Delta T \equiv T_2 T_1$  in terms of  $\Delta p$  and  $\Delta S$ .

### A change of variables

- At this point we need to express  $\mu$  and T as a function of p and S .
- Assume noninteracting Fermi gas to do so:

$$n = \frac{\mu^3}{6\pi^2} + \frac{\mu T^2}{6},$$

$$p = \frac{\epsilon}{3} = \frac{\mu^4}{24\pi^2} + \frac{\mu^2 T^2}{12},$$

$$S = \frac{\pi^2 T}{\mu}.$$

The pressure volume relation expanded:

$$\Delta S \approx \frac{216\pi^6}{\mu_1^{11}T_1} \left(\Delta p\right)^3 - \frac{\omega_1\lambda}{T_1} \frac{36\sqrt{2}\pi^4}{\mu_1^8} (\Delta p)^2 + \dots$$
 Dominates for  $\omega_1 < \frac{\Delta p}{\mu_1^3} 3\sqrt{2}\pi^2$ 

And we are back to nonchiral shockwaves...

For 
$$\omega_1 > \frac{\Delta p}{\mu_1^3} 3\sqrt{2}\pi^2$$
 dominates



$$\Delta S \approx \frac{216\pi^6}{\mu_1^{11}T_1} (\Delta p)^3 - \frac{\omega_1 \lambda}{T_1} \frac{36\sqrt{2}\pi^4}{\mu_1^8} (\Delta p)^2 + \dots$$

And the entropy discontinuity:

$$\Delta S \approx -\frac{\omega_1 \lambda}{T_1} \frac{36\sqrt{2\pi^4}}{\mu_1^8} (\Delta p)^2 + \dots$$

- $\Delta S$  is quadratic in  $(\Delta p)$  in chiral matter instead of being cubic as in nonchiral matter.
- Both rarefaction and compression shockwaves are allowed in chiral matter provided chiral transport dominates!
- Depending on the chirality of fermion, the wave can only propagate either along the vorticity or opposite to the vorticity, but not both.

#### Conclusion

- We find that rarefaction shockwaves are allowed by the second law of thermodynamics in chiral matter.
- Our result is exemplified in a limit  $T/\mu \ll 1$  in a vorticity.
- We expect the qualitative form to hold in other regimes such as that of high temperature and nonzero magnetic field as well.